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Homework #6 Solution

Statistical Mechanics.

1)

Problem 1

Consider an ideal quantum gas of Fermi particles at a temperature T .

- (a) Write the probability $p(n)$ that there are n particles in a given single particle state as a function of the mean occupation number, $\langle n \rangle$.
- (b) Find the root-mean-square fluctuation $\langle (n - \langle n \rangle)^2 \rangle^{1/2}$ in the occupation number of a single particle state as a function of the mean occupation number $\langle n \rangle$. Sketch the result.

Solution

- (a) Let ϵ be the energy of a single particle state, μ be the chemical potential. The partition function is

$$z = \sum_n e^{n(\mu-\epsilon)/kT} = 1 + e^{(\mu-\epsilon)/kT}.$$

The mean occupation number is

$$\langle n \rangle = kT \frac{\partial}{\partial \mu} \ln z = \frac{1}{e^{(\epsilon-\mu)/kT} + 1}.$$

The probability is

$$\begin{aligned} p(n) &= \frac{1}{z} e^{n(\mu-\epsilon)/kT} \\ &= \frac{(1 - \langle n \rangle)^n}{\langle n \rangle^{n-1}}. \end{aligned}$$

- (b) $\langle (n - \langle n \rangle)^2 \rangle = kT \frac{\partial \langle n \rangle}{\partial \mu} = \langle n \rangle (1 - \langle n \rangle).$

So we have $\langle (n - \langle n \rangle)^2 \rangle^{1/2} = \sqrt{\langle n \rangle (1 - \langle n \rangle)}$.

The result is shown in the figure.

Problem 1, cont.

The density of states is

$$D(E) = \frac{4\pi V (2m)^{3/2}}{h^3} E^{1/2}.$$

(a) If $T = 0$, the total number of electrons is

$$N = \int_0^{E_F} D(E) dE = \frac{2}{3} D(E_F) E_F.$$

(b) If $T \neq 0$,

$$N = \int_0^{\infty} \frac{D(E)}{e^{(E-\mu)/kT} + 1} dE.$$

2)

Solution

(a) By the Fermi distribution, the probability for a level ϵ to be occupied is

$$F(\epsilon) = \frac{1}{e^{\beta(\epsilon-\mu)} + 1},$$

so the probability for finding an electron at $\epsilon = \mu + \Delta$ is

$$F(\mu + \Delta) = \frac{1}{e^{\beta\Delta} + 1},$$

and the probability for not finding electrons at $\epsilon = \mu - \Delta$ is given by

$$1 - F(\mu - \Delta) = \frac{1}{e^{\beta\Delta} + 1}.$$

The two probabilities have the same value as required.

(b) When $T > 0$ K, some electrons with $\epsilon < 0$ will be excited to states of $\epsilon > \epsilon_g$. That is to say, vacancies are produced in the some states of $\epsilon < 0$ while some electrons occupy states of $\epsilon > \epsilon_g$. The number of electrons with $\epsilon > \epsilon_g$ is given by

$$\begin{aligned} n_e &= \int_{\epsilon_g}^{\infty} D(\epsilon) \frac{1}{e^{\beta(\epsilon-\mu)} + 1} d\epsilon \\ &= \int_{\epsilon_g}^{\infty} a(\epsilon - \epsilon_g)^{1/2} \frac{1}{e^{\beta(\epsilon-\mu)} + 1} d\epsilon. \end{aligned}$$

The number of vacancies for $\epsilon < 0$ is given by

$$\begin{aligned} n_p &= \int_{-\infty}^0 D(\epsilon) [1 - F(\epsilon)] d\epsilon \\ &= \int_{-\infty}^0 b(-\epsilon)^{1/2} \frac{1}{e^{-\beta(\epsilon-\mu)} + 1} d\epsilon. \end{aligned}$$

By $n_e = n_p$, we have $\mu = \epsilon_g/2$ when $a = b$. We also obtain the equation to determine μ when $a \neq b$,

$$\frac{a}{b} = \frac{e^{\beta(\epsilon+\epsilon_g-\mu)} + 1}{e^{\beta(\epsilon+\mu)} + 1}.$$

For $a > b$, we have

$$\frac{e^{\beta(\epsilon + \epsilon_g - \mu)} + 1}{e^{\beta(\epsilon + \mu)} + 1} > 1,$$

so that $\epsilon + \epsilon_g - \mu > \epsilon + \mu$, i.e., $\mu < \epsilon_g/2$. Hence μ shifts to lower energies.

For $a < b$, $\mu > \epsilon_g/2$, μ shifts to higher energies.

(c) When $T = 0$, by

$$\int_{\epsilon_g}^{\mu} a(\epsilon - \epsilon_g)^{1/2} d\epsilon = n_d,$$

we obtain

$$\mu = \epsilon_g + \left(\frac{3n_d}{2a}\right)^{2/3}.$$

μ shifts to lower energies as T increases.

HW 7, Problem 3 solution:

a) The number of states with energy less than ω is

$$\Sigma(\omega) = \iint_{|\vec{k}| < \omega/c} \frac{d^3r dk}{(2\pi)^3} = \frac{4\pi V}{(2\pi)^3} \int_{|\vec{k}| < \omega/c} k^2 dk = \frac{4\pi V}{(2\pi)^3} \frac{1}{3} \frac{\omega^3}{c^3}$$

Thus, the number of states between ω and $\omega + d\omega$ is

$$g(\omega) = \frac{d\Sigma}{d\omega} = \frac{V}{2\pi^2 c^3} \omega^2$$

b) From $\int_0^{\omega_c} d\omega g(\omega) = 3N$, we get:

$$\omega_c = (18\pi^2 N c^3 / V)^{1/3}$$

c) Since $z=1$, the occupation number is

$$\langle n_{\vec{k}} \rangle = \frac{1}{\exp(\beta \hbar c k) - 1}$$

d) The total energy is:

$$\begin{aligned} \langle E \rangle &= \int_0^{\omega_c} d\omega g(\omega) \langle \varepsilon(\omega) \rangle = \int_0^{\omega_c} d\omega g(\omega) \hbar \omega \left(\frac{1}{2} + \langle n_{\omega} \rangle \right) \\ &= \int_0^{\omega_c} d\omega \frac{V \omega^2}{2\pi^2 c^3} \hbar \omega \left(\frac{1}{2} + \frac{1}{\exp(\beta \hbar \omega) - 1} \right) \\ &= \frac{\hbar V}{2\pi^2 c^3} \int_0^{\omega_c} d\omega \omega^3 \left(\frac{1}{2} + \frac{1}{\exp(\beta \hbar \omega) - 1} \right) \end{aligned}$$

The specific heat is

$$\begin{aligned} C_V &= \left. \frac{dE}{dT} \right|_{V,N} = \frac{\hbar V}{2\pi^2 c^3} \int_0^{\omega_c} d\omega \hbar \omega^4 \frac{\exp(\beta \hbar \omega)}{[\exp(\beta \hbar \omega) - 1]^2} \frac{1}{kT^2} \\ &= \frac{\hbar^2 V}{2\pi^2 c^3 kT^2} \int_0^{\omega_c} d\omega \omega^4 \frac{\exp(\beta \hbar \omega)}{[\exp(\beta \hbar \omega) - 1]^2} \\ &= 9Nk \left(\frac{kT}{\hbar \omega_c} \right)^3 \int_0^{\beta \hbar \omega_c} d(\beta \hbar \omega) (\beta \hbar \omega)^4 \frac{\exp(\beta \hbar \omega)}{[\exp(\beta \hbar \omega) - 1]^2} \\ &= 9Nk \left(\frac{kT}{\hbar \omega_c} \right)^3 \int_0^{\beta \hbar \omega_c} dx \frac{x^4 \exp(x)}{[\exp(x) - 1]^2} \end{aligned}$$

For small T, $\beta \hbar \omega \rightarrow \infty$, the integration gives a constant so $C_V \sim T^3$.

For large T, $\beta \hbar \omega \rightarrow 0$, $\exp(x) \sim 1+x$, the integration gives $(\beta \hbar \omega_c)^3 / 3$.

So $C_V \sim 3Nk$ as in classical gas.

Determine the form of $\rho(\omega)$, the spectral density of the energy.

(B) What is the temperature dependence of the energy \bar{E} ?

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Solution:

(A) The density of states is $8\pi V p^2 dp/h^3$, or $V \omega^2 d\omega/\pi^2 c^3$. Then the number of photons is

$$\begin{aligned} \bar{N} &= \int \frac{V}{\pi^2 c^3} \omega^2 \frac{1}{e^{h\omega/kT} - 1} d\omega \\ &= \frac{V}{\pi^2 c^3} \left(\frac{kT}{h}\right)^3 \int_0^\infty \frac{x^2 dx}{e^x - 1} \propto T^3. \end{aligned}$$

$$\begin{aligned} \text{(b), (C)} \quad \frac{\bar{E}}{V} &= \int \frac{\omega^2}{\pi^2 c^3} \frac{h\omega}{e^{h\omega/kT} - 1} d\omega \\ &= \frac{(kT)^4}{\pi^2 c^3 h^3} \int \frac{\xi^3 d\xi}{e^\xi - 1}. \end{aligned}$$

Hence

$$\rho(\omega) = \frac{h}{\pi^2 c^3} \frac{\omega^3}{e^{h\omega/kT} - 1},$$

and $\bar{E} \propto T^4$.

If $\mu = 0$, the above occurs:

Problem 4

Consider a photon gas enclosed in a volume V and in equilibrium at temperature T . The photon is a massless particle, so that $\epsilon = pc$.

(a) What is the chemical potential of the gas? Explain.

(B) Determine how the number of photons in the volume depends upon the temperature.

(c) One may write the energy density in the form

$$\frac{\bar{E}}{V} = \int_0^\infty \rho(\omega) d\omega.$$

$$\rho(\omega) = \frac{4\pi}{h^3 c^3} \omega^3 \frac{1}{e^{h\omega/kT} - 1}$$

Problem 4.1
The density of state is $g(\epsilon) = \alpha V \epsilon^2$

α is a constant.

Grand thermodynamic potential

$$\begin{aligned}\Omega(T, V, \mu) &= -k_B T \ln \mathcal{Z} \\ &= +k_B T \int d\epsilon g(\epsilon) \ln(1 - e^{-\beta\epsilon}) \\ &= k_B T \alpha V \int d\epsilon \epsilon^2 \ln(1 - e^{-\beta\epsilon}) \\ &= k_B T \alpha V \int d\left(\frac{\epsilon^3}{3}\right) \ln(1 - e^{-\beta\epsilon}) \\ &= -k_B T \alpha V \int d\epsilon \frac{\beta e^{-\beta\epsilon}}{1 - e^{-\beta\epsilon}} \cdot \frac{\epsilon^3}{3} \\ &= -\cancel{\alpha V} \int d\epsilon \frac{\epsilon^3}{3} \frac{1}{e^{\beta\epsilon} - 1}\end{aligned}$$

$$\begin{aligned}P &= -\frac{\partial \Omega}{\partial V} \Big|_{T, \mu} = +\alpha \int d\epsilon \frac{\epsilon^3}{3} \frac{1}{e^{\beta\epsilon} - 1} \\ &= \frac{1}{3V} \int d\epsilon g(\epsilon) \frac{\epsilon}{e^{\beta\epsilon} - 1} \\ &= \frac{1}{3V} \cdot \langle E \rangle\end{aligned}$$

$$P = \frac{\cancel{2}}{3V} \langle E \rangle$$

$$\text{since } \langle E \rangle \sim T^4, \quad P \sim T^4.$$

Homework 7. # Problem 5.

$$Z = \sum_{N,i} e^{-\beta(E_i - \mu N)}$$

$$Z = e^{\beta \mu} : \text{fugacity}$$

$$\left. \frac{\partial \ln Z}{\partial \beta} \right|_{\mu, V} = \frac{1}{Z} \left. \frac{\partial Z}{\partial \beta} \right|_{\mu, V}$$

$$= \frac{1}{Z} \left. \frac{\partial}{\partial \beta} \left(\sum_{N,i} e^{-\beta(E_i - \mu N)} \right) \right|_{\mu, V}$$

$$= \frac{1}{Z} \sum (\mu N - E_i) e^{-\beta(E_i - \mu N)}$$

$$\left. - \frac{\partial \ln Z}{\partial \beta} \right|_{\mu, V} = E$$

$$\left. - \frac{\partial \ln Z}{\partial \beta} \right|_{\mu, V} = \frac{1}{Z} \sum (E_i - \mu N) e^{-\beta(E_i - \mu N)}$$

$$= E - \mu \langle N \rangle$$

$$E = - \left. \frac{\partial \ln Z}{\partial \beta} \right|_{\mu, V} + \mu \langle N \rangle$$