

Ideal Fermi gas

$$\text{For } T \rightarrow 0 \quad \langle n_{\mathbf{k}} \rangle = \frac{1}{e^{\beta(\epsilon_{\mathbf{k}} - \mu)} + 1} = \begin{cases} 1 & \text{if } \epsilon_{\mathbf{k}} < \mu \\ 0 & \text{if } \epsilon_{\mathbf{k}} > \mu \end{cases}$$

this limiting value $\mu(T \rightarrow 0)$ is called the Fermi energy ϵ_F . All one-particle states with $\epsilon_{\mathbf{k}} < \epsilon_F$ is occupied at $T=0$. For ideal gas $\epsilon_{\mathbf{k}} = \hbar^2 k^2 / 2m$

⇒ There's a corresponding Fermi wave vector k_F .

Since
$$N = g \sum_{\vec{k}} \langle n_{\vec{k}} \rangle = gV \int_0^{k_F} \frac{d^3k}{(2\pi)^3} = \frac{gV k_F^3}{6\pi^2}$$

⇒
$$k_F = \left(\frac{6\pi^2 n}{g} \right)^{1/3} \quad \epsilon_F = \frac{\hbar^2 k_F^2}{2m} = \frac{\hbar^2}{2m} \left(\frac{6\pi^2 n}{g} \right)^{2/3}$$

$n = N/V$

(Note classical gas has large # of microstate at $T=0$, $\Omega = V^N/N!$, but Fermi gas has only one unique ground state, $\Omega = 1$)

Low temperature expansion: $k_B T \ll \epsilon_F$ ($\ln z = |\beta \mu| \gg 1$)

$$n = g \int_0^\infty \frac{d^3k}{(2\pi)^3} \frac{1}{z^{-1} \exp\left(\frac{\beta \hbar^2 k^2}{2m}\right) + 1}$$

$$= \frac{g}{\lambda^3} \frac{2}{\sqrt{\pi}} \int_0^\infty \frac{dx \cdot x^{1/2}}{z^{-1} \exp(x) + 1}$$

$$\approx \frac{g}{\lambda^3} \frac{(\ln z)^{3/2}}{(3/2)!} \left[1 + \frac{\pi^2}{6} \frac{3}{4} (\ln z)^{-2} + \dots \right]$$

Solve for $\ln z$ gives

$$\ln z \approx \beta \epsilon_F \left[1 - \frac{\pi^2}{12} \left(\frac{k_B T}{\epsilon_F} \right)^2 \right]$$

Similarly the pressure

$$P \approx P_F \left[1 + \frac{5}{12} \pi^2 \left(\frac{k_B T}{\epsilon_F} \right)^2 \right] \quad P_F = \frac{2}{5} n \epsilon_F$$

and energy

$$E = \frac{3}{2} PV = \frac{3}{2} P_F V \left[1 + \frac{5}{12} \pi^2 \left(\frac{kT}{\epsilon_F} \right)^2 \right]$$

Specific heat

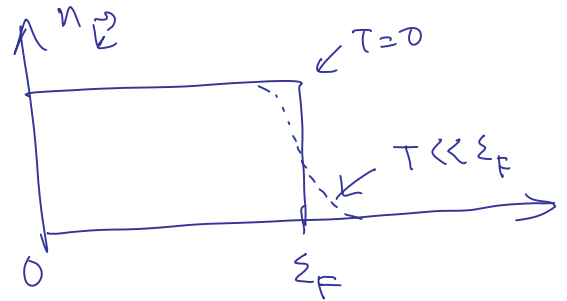
$$C_V = \frac{\partial E}{\partial T} \Big|_V \approx \frac{\pi^2}{2} N k_B \frac{kT}{\epsilon_F} \sim T$$

* μ (or $\ln z$) goes from positive ϵ_F at $T \rightarrow 0$ to negative value at large T .

* $C_V \sim T$ at low T because only a small # of particles ($\sim N \frac{kT}{\epsilon_F}$) are excited. each can absorb energy kT , so

$$\Delta E \sim kT \cdot N \frac{kT}{\epsilon_F}$$

$$\Rightarrow C = \frac{\Delta E}{T} \approx Nk \cdot \frac{kT}{\epsilon_F}$$



Ideal Bose gas

$$\langle n_k \rangle_+ = \frac{1}{e^{\beta(\epsilon_k - \mu)} - 1}$$

* $\langle n_k \rangle$ must be positive so $\mu < \epsilon_k$ for all k .

Since the smallest ϵ_k is at $k=0 \Rightarrow \mu < 0$

* $T \rightarrow 0$ $n_k \rightarrow 0$ except $k=0$, $\epsilon_0 = \mu = 0$, all boson occupy the lowest energy level. Due to this fact, at finite T , special care has to

taken when convert $\sum_{\vec{k}}$ to integration.

Specifically, the sum over microstates would be

$$\sum_{\vec{k}} \sum_{\vec{k}} \rightarrow V \int \frac{d^3 k}{(2\pi)^3} = \int_0^\infty g(\epsilon) d\epsilon$$

where the density of states $g(\epsilon) = \frac{2\pi V}{h^3} (2m)^{3/2} \epsilon^{1/2}$

goes to zero for $\epsilon = 0$. In other words, the above integral approximation to $\sum_{\vec{k}}$ is very bad. To fix this, we explicitly account for $\vec{k} = 0$ term

$$\sum_{\vec{k}} \rightarrow (\vec{k} = 0) + \int \frac{d^3 k}{(2\pi)^3}$$

The grand potential, thus, is

$$\Omega(T, V, \mu) = -kT \ln \mathcal{Z} = kT \sum_{\vec{k}} \ln(1 - z e^{-\beta \epsilon})$$

$$= kT \ln(1 - z) + V \int \frac{d^3 k}{(2\pi)^3} \ln(1 - z e^{-\beta \epsilon}) \cdot kT$$

$$= kT \ln(1 - z) - \frac{2}{3} kT \int_0^\infty d\epsilon \frac{g(\epsilon) \cdot \epsilon}{z^{-1} e^{\beta \epsilon} - 1}$$

and

$$N = \sum_{\vec{k}} n_{\vec{k}} = \sum_{\vec{k}} \frac{1}{z^{-1} e^{\beta \epsilon_{\vec{k}}} - 1} = \frac{1}{z^{-1} - 1} + \int_0^\infty d\epsilon \frac{g(\epsilon)}{z^{-1} e^{\beta \epsilon} - 1}$$

$$\equiv N_0 + N_e$$

N_0 : # of bosons in ground state

N_e : # of bosons in excited state.

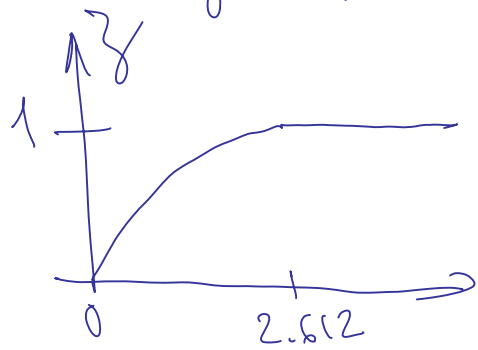
Define a function $g_n(z) = \frac{1}{(n-1)!} \int_0^\infty \frac{x^{n-1} dx}{z^{-1} \exp(x) - 1}$

$$\Omega = kT \ln(1-z) - \frac{V}{\lambda^3} g_{5/2}(z) kT$$

$$N = \frac{z}{1-z} + \frac{V}{\lambda^3} g_{3/2}(z)$$

if one plots $g_{3/2}(z)$ as function of z , one sees that it has a maximum at $z=1$ and $N_e^{\max} = \frac{V}{\lambda^3} g_{3/2}(1) \cong (V/\lambda^3) 2.612 \sim V \cdot T^{3/2}$

at very small T , $N_e^{\max} \ll N$, $z=1$ and $N \sim N_0$
at very large T , $N_e^{\max} \gg 1$, $z < 0$ and $N_0 = 0$



\Rightarrow there's a critical temperature T_c where $N_e^{\max} = N$

$$\Rightarrow T_c = \left(\frac{N}{V}\right)^{2/3} \frac{h^2}{2\pi m (g_{3/2}(1))^{2/3}}$$

$T < T_c$, $z=1$, $N_0 \geq 0$. There's a macroscopic occupation of single-particle state $\sum_{k=0}^{\infty}$. This is Bose-Einstein condensation.

$T > T_c$, $N_0 = 0$ and z is obtained from $N_e(z) = N$
or $\frac{N\lambda^3}{V} = g_{3/2}(z)$

Pressure

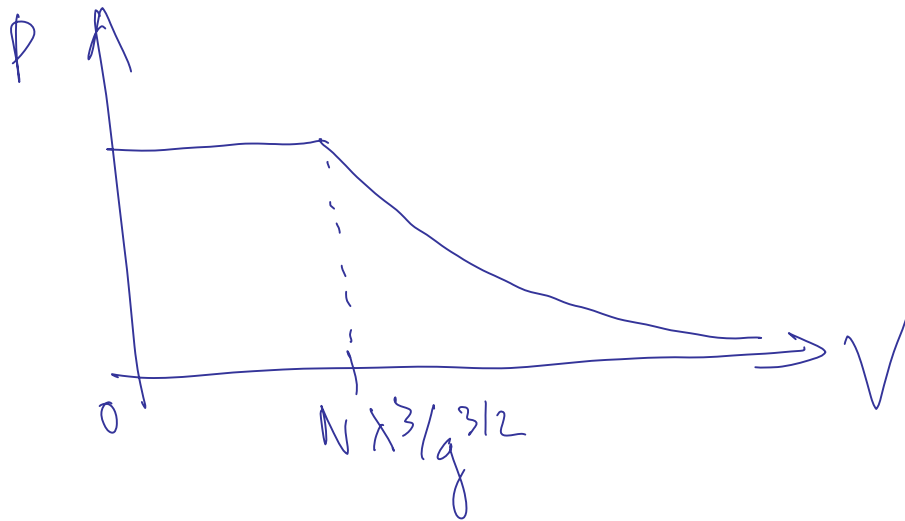
the particles with $\vec{v} = 0$ do not contribute

$$P = - \frac{\partial \Omega}{\partial V} = kT \frac{g_{5/2}(z)}{\lambda^3}$$

$$T < T_c, \quad z = 1 \quad P = \frac{g_{5/2}(1)}{\lambda^3} \cdot kT \quad \text{independent of } V \text{ or } N$$

$$T \rightarrow \infty, \quad g_{5/2}(z) \approx g_{3/2}(z) \approx z$$

$$\text{So } P = kTN/V \Rightarrow \text{classical ideal gas}$$



Specific heat:

$$\bar{E} = - \frac{\partial \ln \Xi}{\partial \beta} \Big|_{z, V} = \frac{3}{2} kT \frac{V}{\lambda^3} g_{5/2}(z) = \frac{3}{2} PV$$

$$C_V = \frac{\partial \bar{E}}{\partial T} \Big|_{V, N}$$

$$T < T_c: \quad C_V = \frac{15}{4} g_{5/2}(1) \cdot \frac{Vk}{\lambda^3} \sim T^{3/2}$$

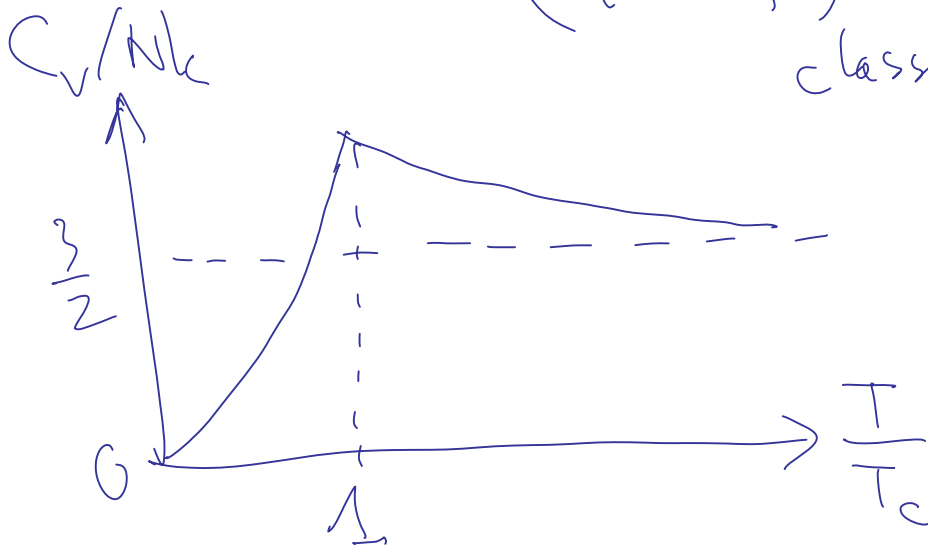
$$T > T_c: C_v = \left(\frac{15}{4} \frac{g_{5/2}(z)}{g_{3/2}(z)} - \frac{9}{4} \frac{g_{3/2}(z)}{g_{1/2}(z)} \right) Nk$$

(See textbook page 323 for derivation)

for $T \rightarrow \infty$, $g_{5/2}(z) \approx g_{3/2}(z) \approx z$

$$C_v = \left(\frac{15}{4} - \frac{9}{4} \right) Nk = \frac{3}{2} Nk$$

classical ideal gas



$C_v \sim T^{3/2}$ because at finite T , there's an occupation of states upto k_{\max} such that $\frac{\hbar^2 k_{\max}^2}{2m} \sim Tk$

the number of states is $\frac{k_{\max}^3}{(2\pi)^3}$. So $E \sim V \cdot \frac{k_{\max}^3}{(2\pi)^3} \cdot kT \sim T^{5/2}$

$\Rightarrow C_v \sim T^{3/2}$

This is in contrast to Fermion $C_v \sim T$ because only particles near Fermi level are excited $E \sim V \frac{k_F^2}{(2\pi)^2} \cdot \frac{kT}{\epsilon_F} \cdot kT$